

How Hall electric fields intrinsically chaotize and heat ions during collisionless magnetic reconnection

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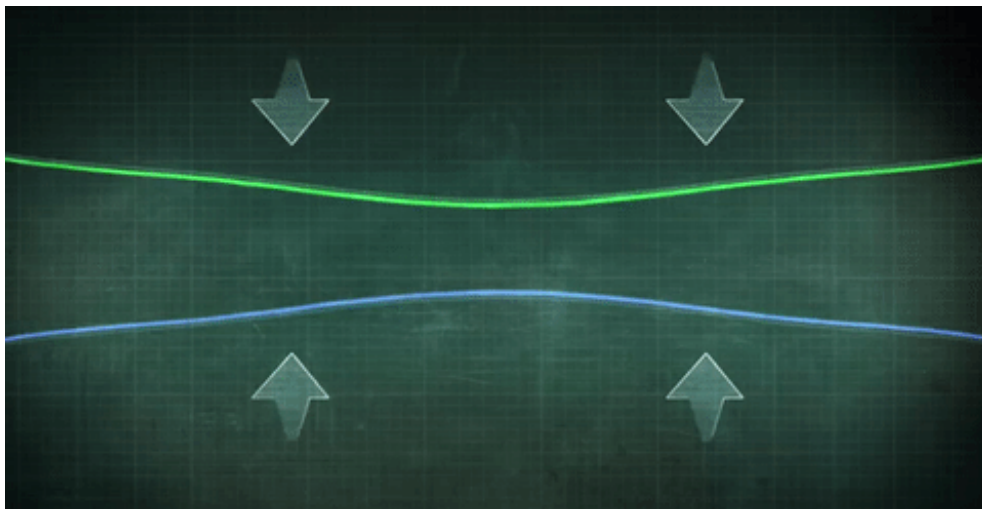
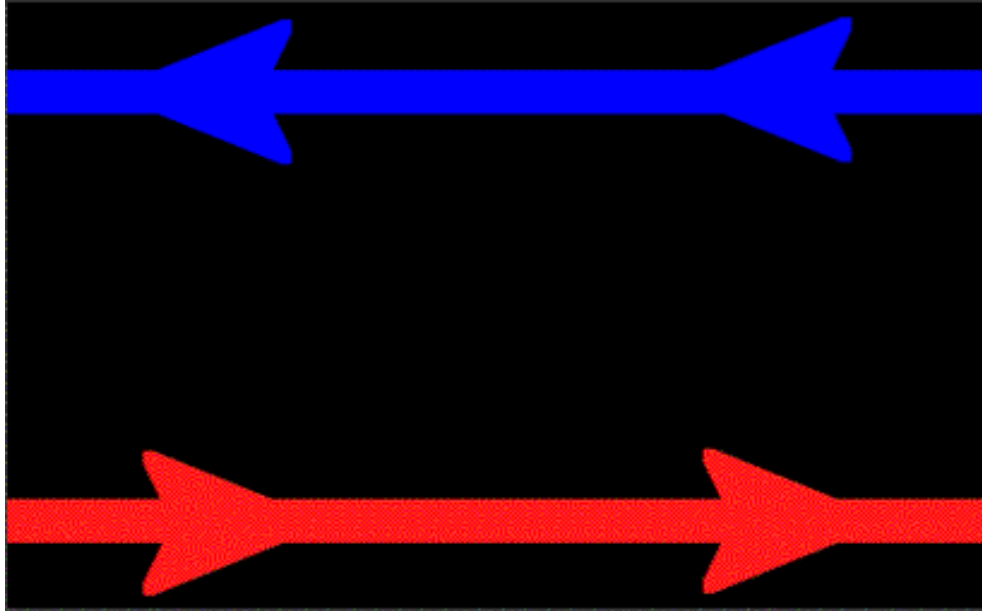


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Outline

- Magnetic reconnection and ion heating
- Stochastic heating
- Fluid derivation using canonical vorticity
- Kinetic analysis
- Particle-in-cell simulations
- Conclusion

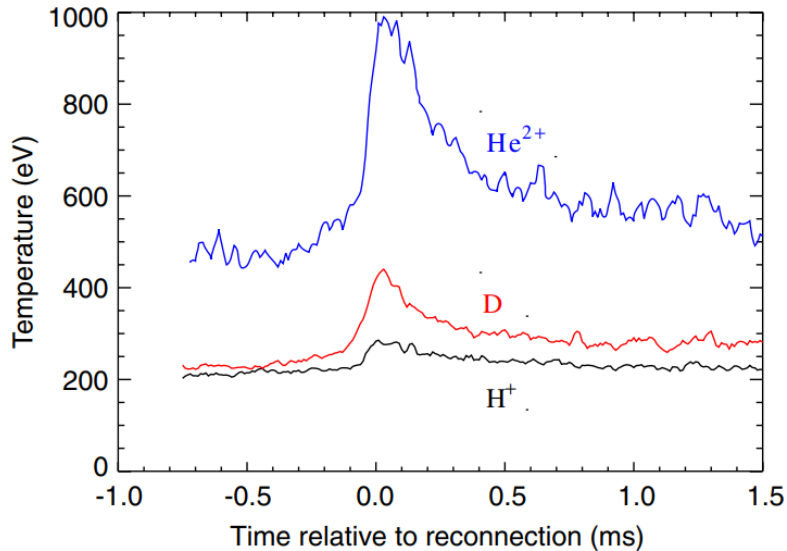
Magnetic Reconnection



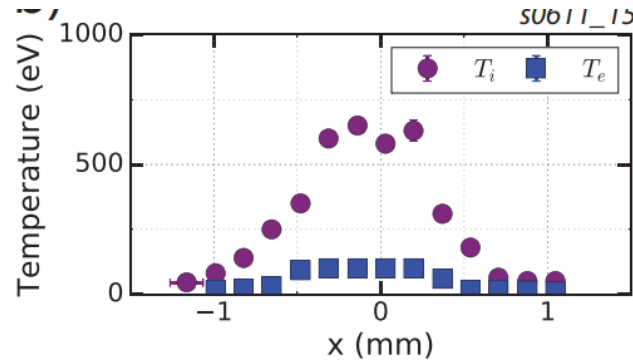
- Fundamental magnetized plasma phenomenon
- Opposing magnetic field lines come together, annihilate, and reconnect
- Common in magnetized plasmas; often abrupt and spontaneous
- Stored magnetic energy \Rightarrow particle kinetic/thermal energy

Ion Heating in Reconnection

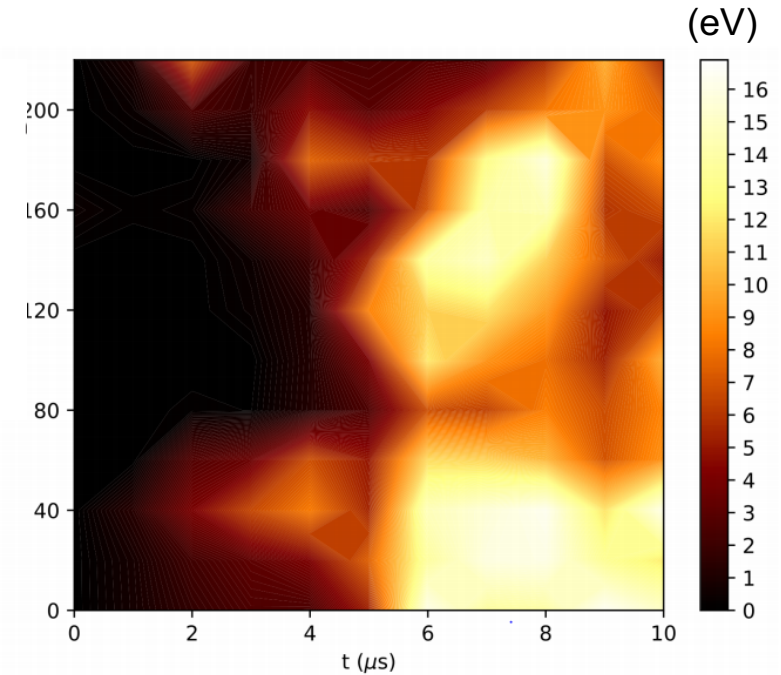
- Fast, strong ion heating is observed with reconnection



Fiksel et al, 2009



Hare et al, 2017



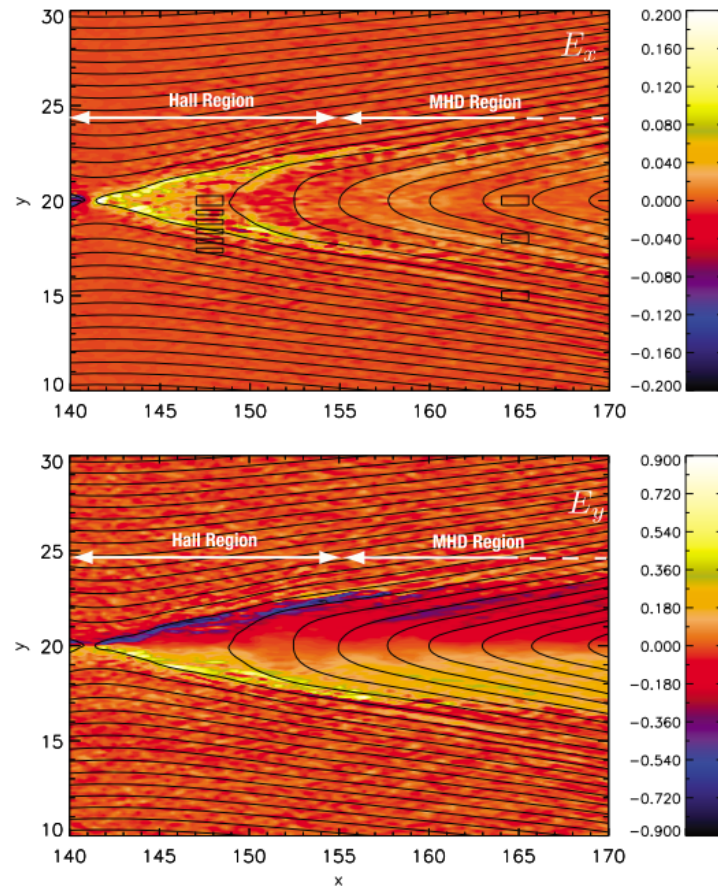
Seo et al., 2020

- Mass-dependent (higher mass => higher temperature)
- Related to coronal heating problem

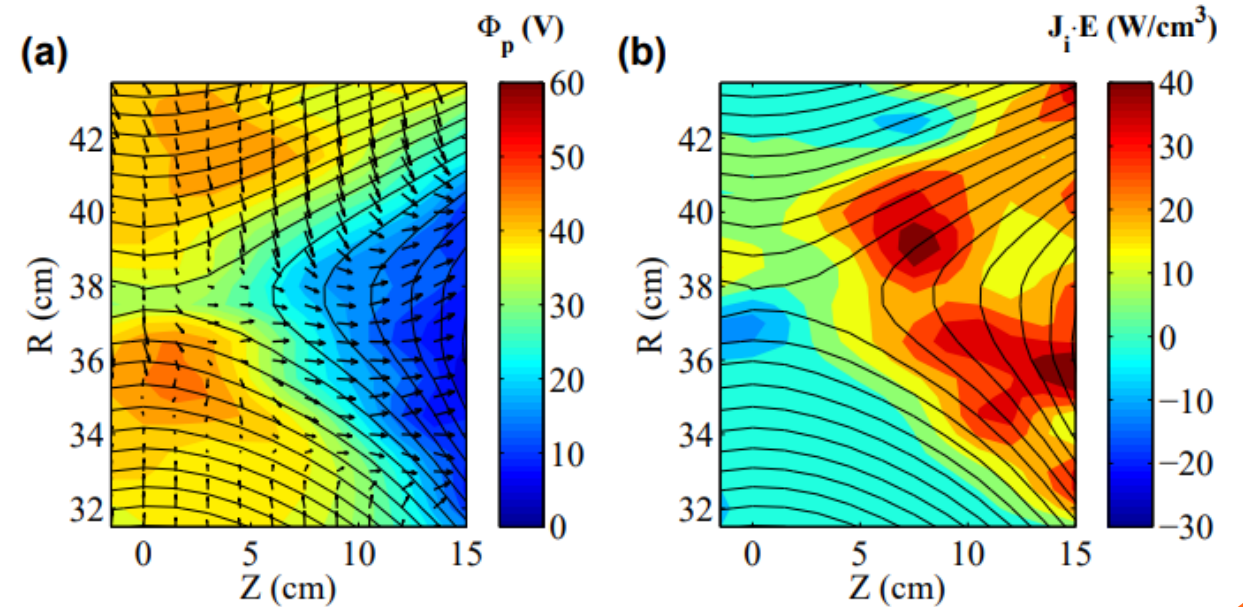
Ion Heating in Reconnection

- In-plane Hall electric fields are attributed to ion energization
- No consensus on the heating mechanism

Simulation (Aunai et al., 2011)



Experiment (Yoo et al., 2013)

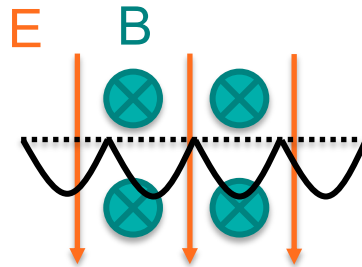


Stochastic Heating

- Occurs when the potential gradient perpendicular to the local magnetic field is large enough

$$\frac{m_i}{q_i B^2} |\nabla_{\perp}^2 \phi| > 1$$

- Chaotic motion => significant heating
- Involves breakdown of guiding center motion

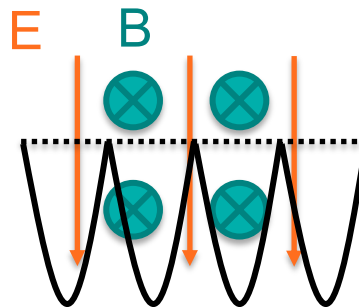


Stochastic Heating

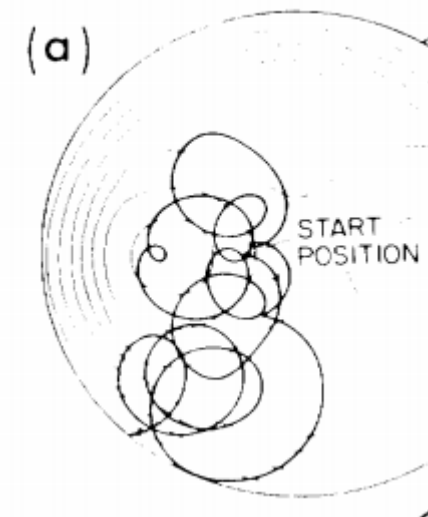
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- Observed in the solar wind and corona, collisionless shocks, and experiments (Vranjes et al., 2009; 2010; Chandran et al., 2010; McChesney et al., 1987; Stasiewicz et al., 2020)



McChesney et al., 1987

Stochastic Heating

- Consider

$$\mathbf{E} = E \hat{y} \sin(ky - \omega t),$$

$$\mathbf{B} = B \hat{z}.$$

- Then,

$$\mathbf{v}_E = \frac{\mathbf{E} \times \mathbf{B}}{B^2} = \hat{x} \frac{E}{B} \sin(ky - \omega t)$$

- and

$$\begin{aligned} \mathbf{v}_p &= \frac{m}{qB^2} \frac{d\mathbf{E}_\perp}{dt}, \\ &= \hat{y} \frac{mE}{qB^2} \frac{d}{dt} \sin(ky - \omega t), \\ &= \hat{y} \frac{mE}{qB^2} \left(k \frac{dy}{dt} - \omega \right) \cos(ky - \omega t). \end{aligned}$$

- Solution is

$$\mathbf{v}_p = -\hat{y} \frac{\omega m E}{q B^2} \frac{\cos(ky - \omega t)}{1 - \alpha \cos(ky - \omega t)},$$

where

$$\alpha = \frac{mkE}{qB^2} \iff \frac{m_i}{q_i B^2} |\nabla_\perp^2 \phi| > 1$$

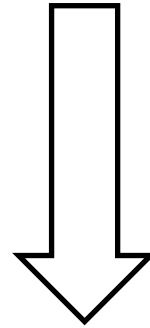
- When $\alpha > 1$, polarization drift \gg ExB drift \Rightarrow nonlinear chaotic motion
- Phase space broadening \Rightarrow (perpendicular) heating

Canonical Vorticity

$$m_e \frac{D\mathbf{u}_e}{Dt} = q_e \mathbf{E} + q_e \mathbf{u}_e \times \mathbf{B}$$

$$m_e \frac{\partial \mathbf{u}_e}{\partial t} + m_e \mathbf{u}_e \cdot \nabla \mathbf{u}_e = q_e \mathbf{E} + q_e \mathbf{u}_e \times \mathbf{B}$$

Take curl



$$\frac{\partial \mathbf{Q}_e}{\partial t} = \nabla \times (\mathbf{u}_e \times \mathbf{Q}_e)$$

Canonical Vorticity

$$\begin{aligned}\mathbf{Q}_e &= m_e \nabla \times \mathbf{u}_e + q_e \mathbf{B} \\ &= \nabla \times (m_e \mathbf{u}_e + q_e \mathbf{A}) \\ &= \nabla \times \mathbf{P}_e\end{aligned}$$

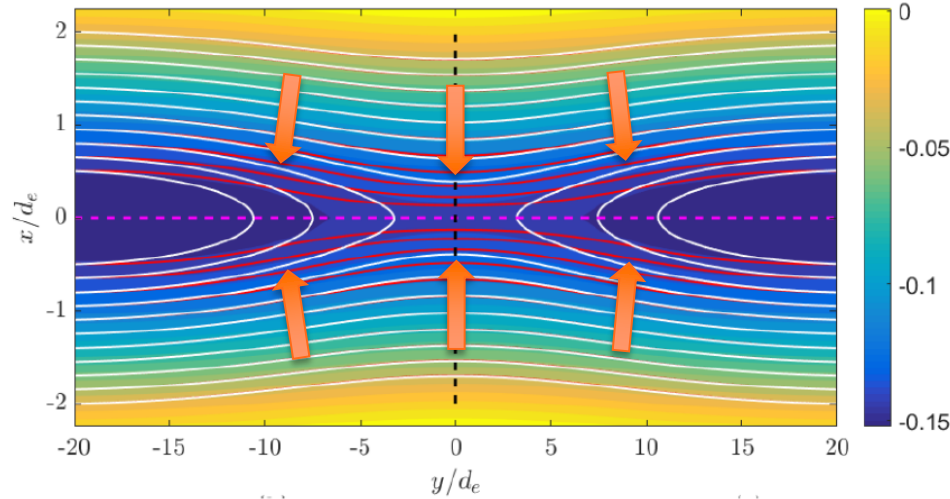
Q is curl of canonical momentum, or **canonical vorticity**

$$\frac{\partial \mathbf{Q}_e}{\partial t} = \nabla \times (\mathbf{u}_e \times \mathbf{Q}_e) \quad \longleftrightarrow \quad \frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{U} \times \mathbf{B})$$

Q is frozen into the electron fluid

Q flux is conserved => **Q cannot reconnect**

In-plane Potential



- White: \mathbf{B}
- Red: \mathbf{Q}
- Color: ϕ
- Arrows: \mathbf{E}

$$q_e \mathbf{E} = m_e \frac{D\mathbf{u}_e}{Dt} - q_e \mathbf{u}_e \times \mathbf{B} + \frac{\nabla p_e}{n_e}$$

$$q_e \mathbf{E} = m_e \frac{\partial \mathbf{u}_e}{\partial t} - \mathbf{u}_e \times \mathbf{Q}_e + \nabla \left(\frac{m_e u_e^2}{2} \right) + \frac{\nabla p_e}{n_e}$$

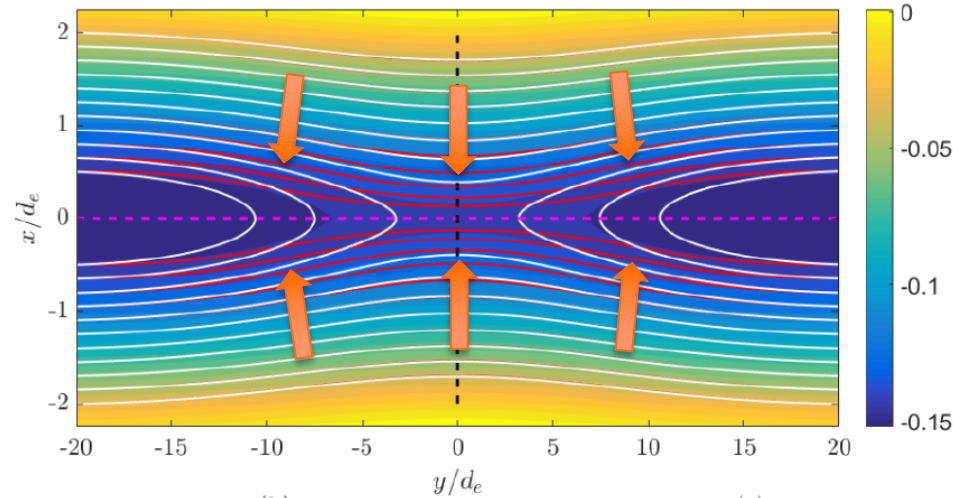
$$q_e E_y = m_e \frac{\partial u_{ey}}{\partial t} - u_{ex} Q_{ex} + u_{ex} Q_{ez} + \frac{\partial}{\partial y} \left(\frac{m_e u_e^2}{2} \right)$$

$$E_y = -\frac{\partial}{\partial y} \left(\frac{m_e [u_e^2 + u_{ey}^2]}{2e} \right)$$

Similarly,

$$E_x = -\frac{\partial}{\partial x} \left(\frac{B^2}{2\mu_0 n_e e} \right)$$

In-plane Potential



- White: \mathbf{B}
- Red: \mathbf{Q}
- Color: ϕ
- Arrows: \mathbf{E}

Stochastic ion heating
criterion:

$$\frac{m_i}{q_i B^2} |\nabla_{\perp}^2 \phi| > 1 \quad \Rightarrow$$

- Inflow criterion $L_x^2 < d_i^2$
- Outflow criterion $m_i / m_e > 1$

Stochastic heating is intrinsic to collisionless reconnection!

Kinetic Analysis

- Harris equilibrium magnetic field (exact kinetic equilibrium)

$$B_y(x) = 2\sqrt{\mu_0 n_0 k_B T} \tanh\left(\frac{x}{\lambda}\right),$$

- Associated electric field:

$$E_x(x) = -VB_y(x) = -2V\sqrt{\mu_0 n_0 k_B T} \tanh\left(\frac{x}{\lambda}\right),$$

- Insert into stochastic heating criterion:

$$\frac{m_i}{q_i [B_y^2(x) + B_z^2]} \left| \frac{\partial E_x(x)}{\partial x} \right| > 1,$$

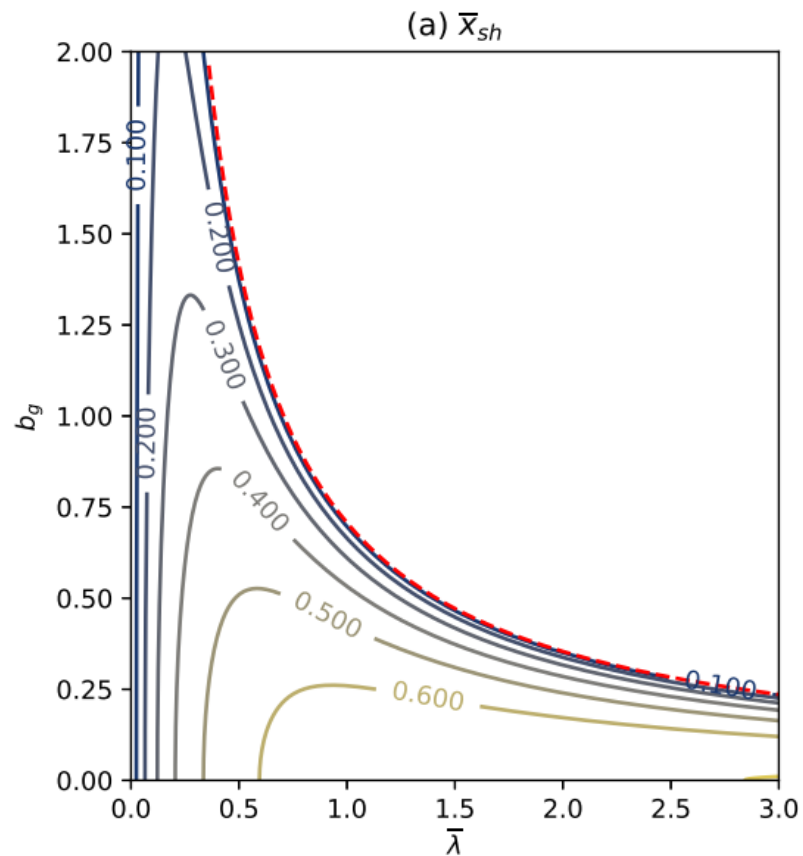
- Equation to solve (b_g is the relative guide field strength)

$$2\frac{\lambda^2}{d_i^2} \cosh^2\left(\frac{x}{\lambda}\right) \left[\tanh^2\left(\frac{x}{\lambda}\right) + b_g^2 \right] < 1,$$

Kinetic Analysis

- Solution is (bar means normalized to d_i)

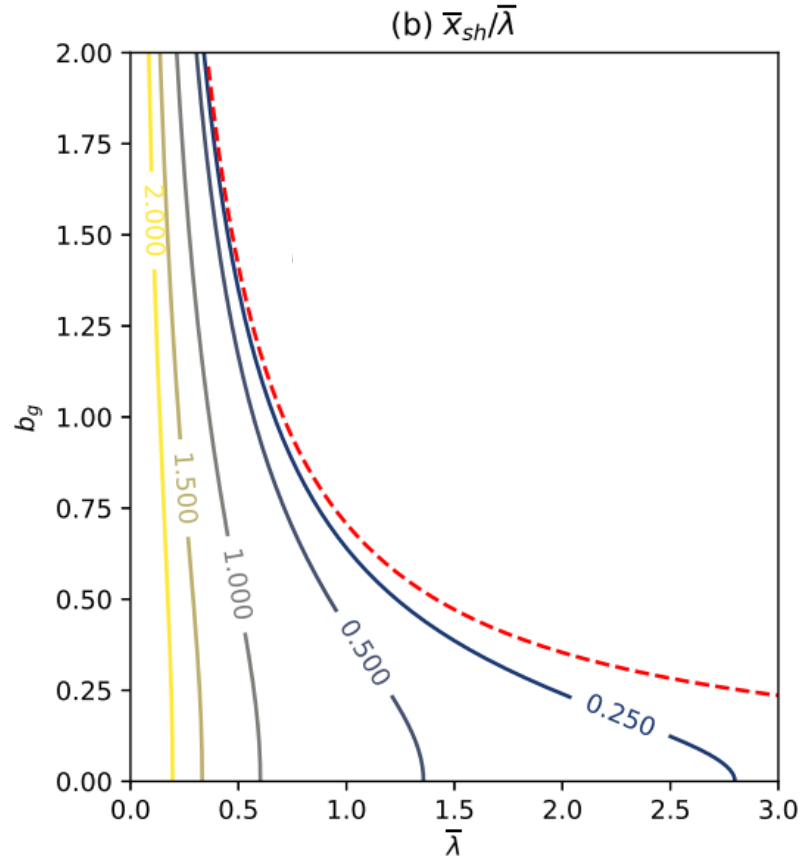
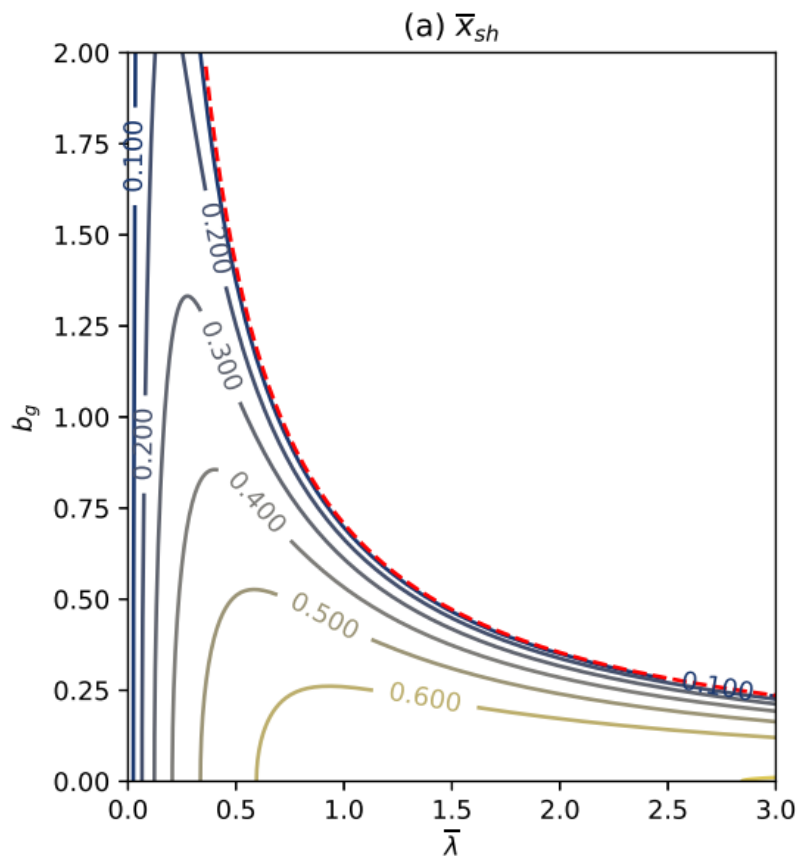
$$\bar{x}_{sh} = \frac{\bar{\lambda}}{2} \ln \frac{\bar{\lambda}^2(1 - b_g^2) + 1 + \sqrt{(1 - 2\bar{\lambda}^2 b_g^2)(2\bar{\lambda}^2 + 1)}}{\bar{\lambda}^2(1 + b_g^2)}.$$



Kinetic Analysis

- Solution is (bar means normalized to d_i)

$$\bar{x}_{sh} = \frac{\bar{\lambda}}{2} \ln \frac{\bar{\lambda}^2(1 - b_g^2) + 1 + \sqrt{(1 - 2\bar{\lambda}^2 b_g^2)(2\bar{\lambda}^2 + 1)}}{\bar{\lambda}^2(1 + b_g^2)}$$



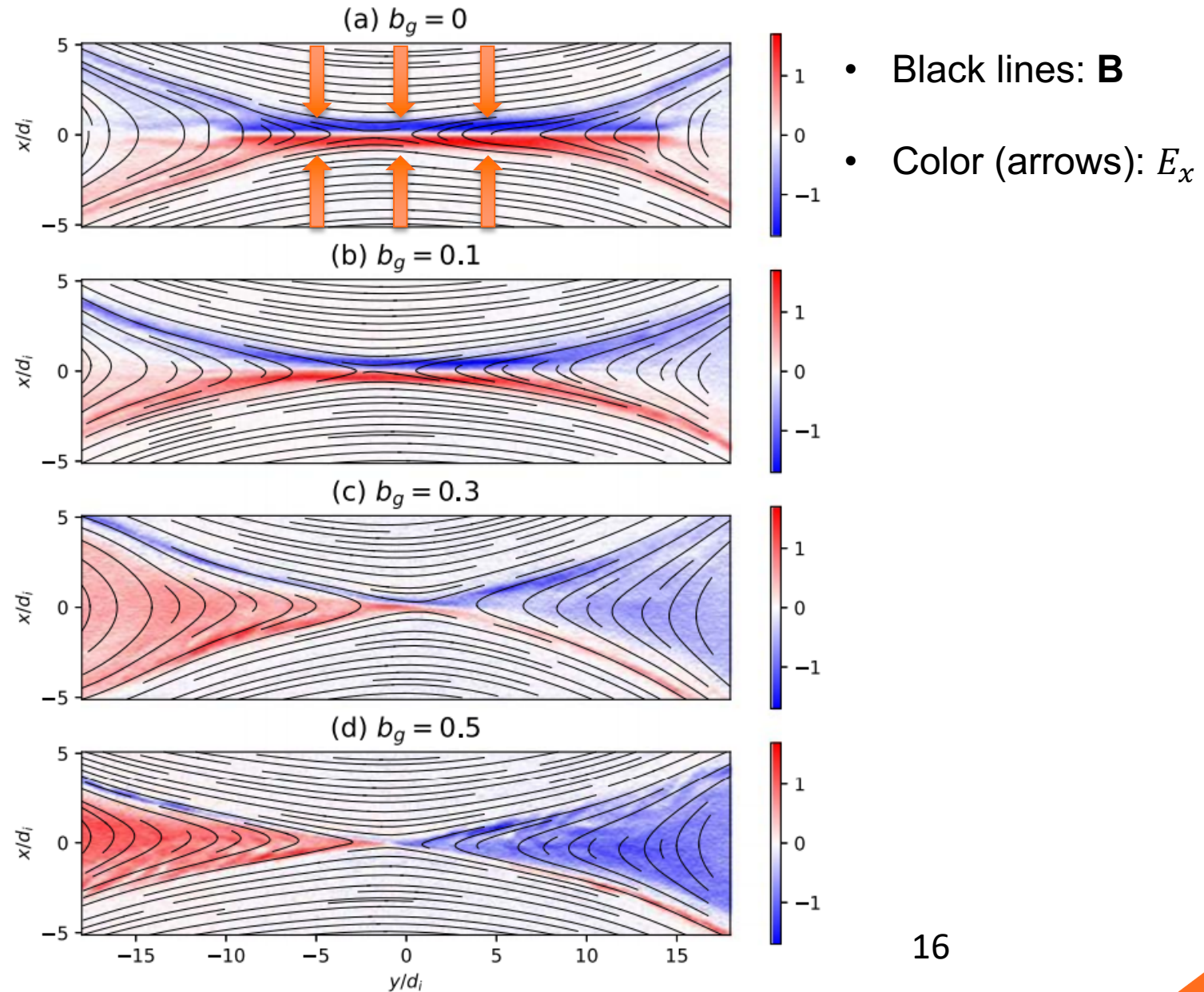
Conclusions:

1. Heating extent \downarrow as
guide field \uparrow
sheet thickness \uparrow
2. Above a threshold guide field,
NO stochastic heating

Particle-in-cell (PIC) Simulation

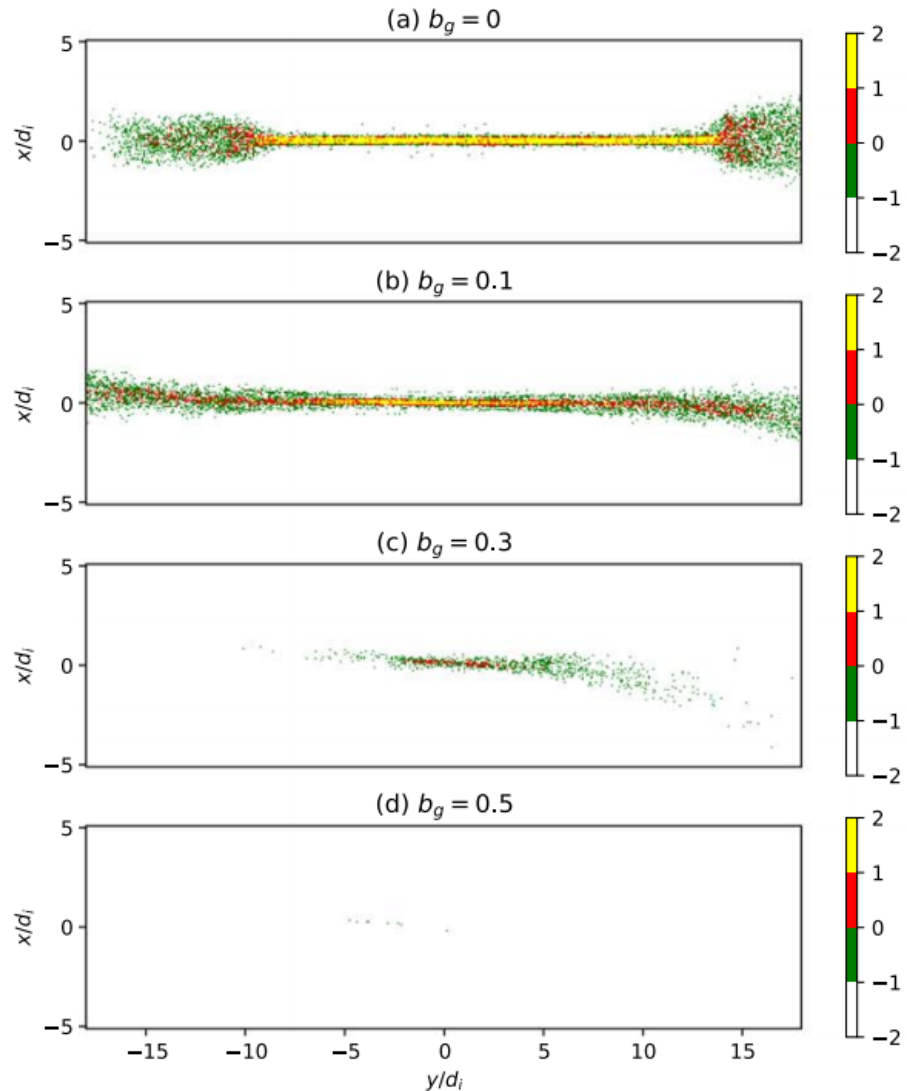
SMILEI code
(Derouillat et al., 2018)

$$m_i/m_e = 100$$

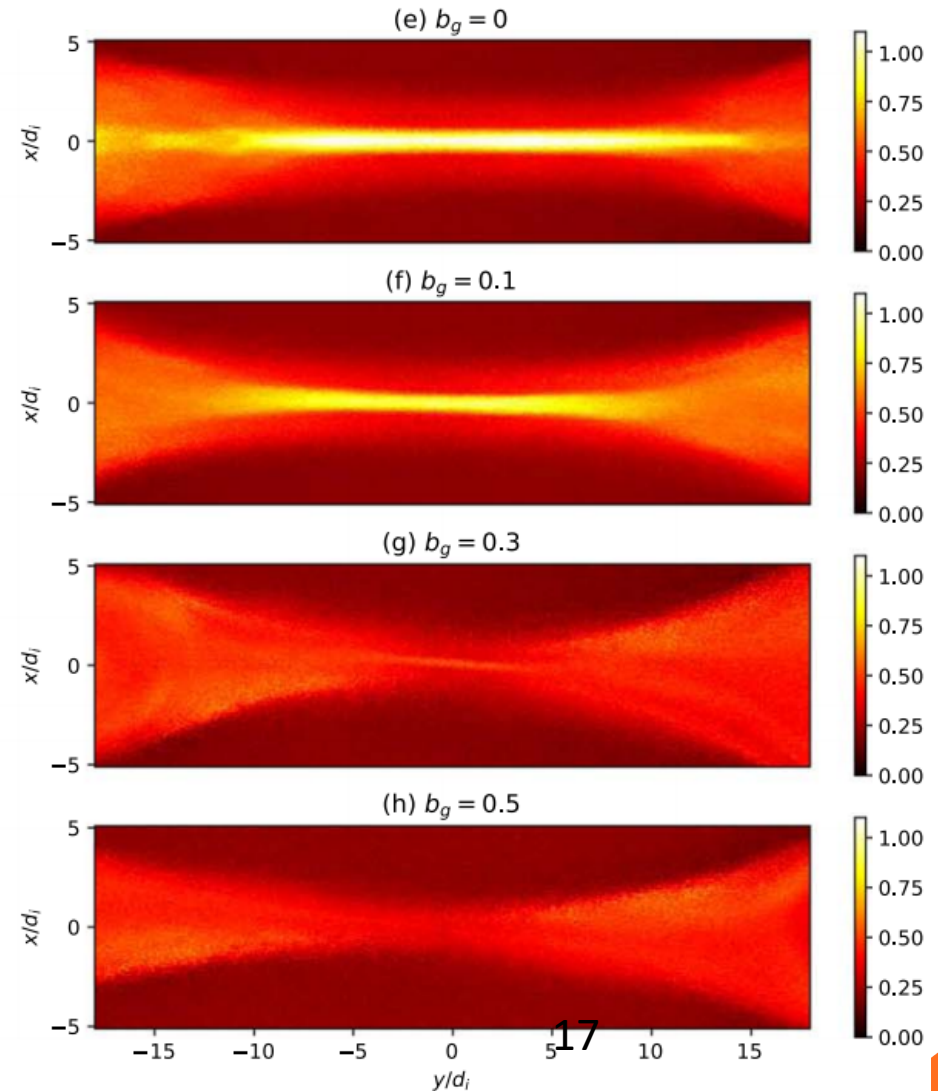


Particle-in-cell (PIC) Simulation

- Stochastic Heating Criterion



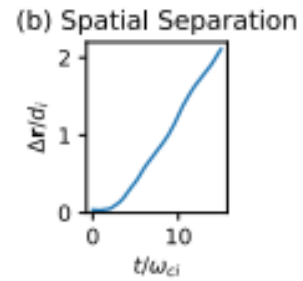
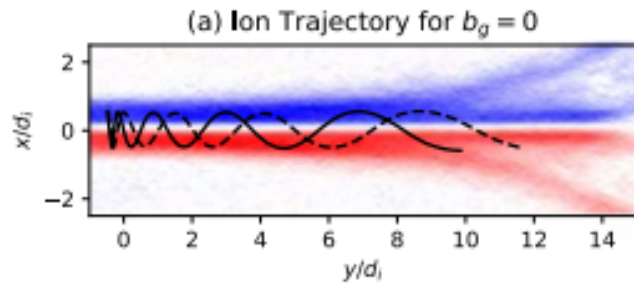
- Ion Temperature



Test-Particle Simulation

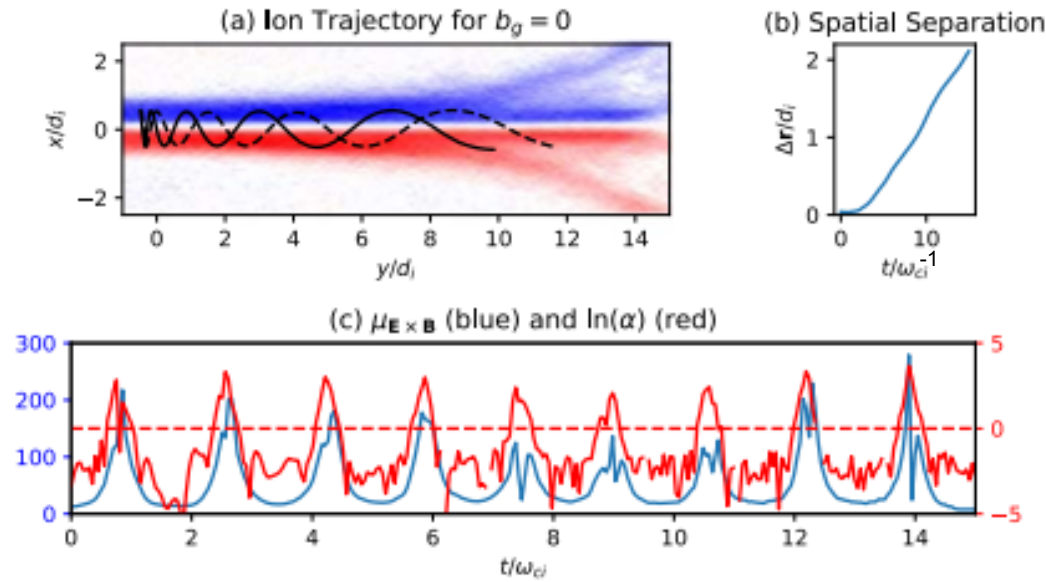
- Two indicators of chaotic behavior:
 1. Spatial divergence of initially close particles
 2. Violation of magnetic moment μ

Test-Particle Simulation



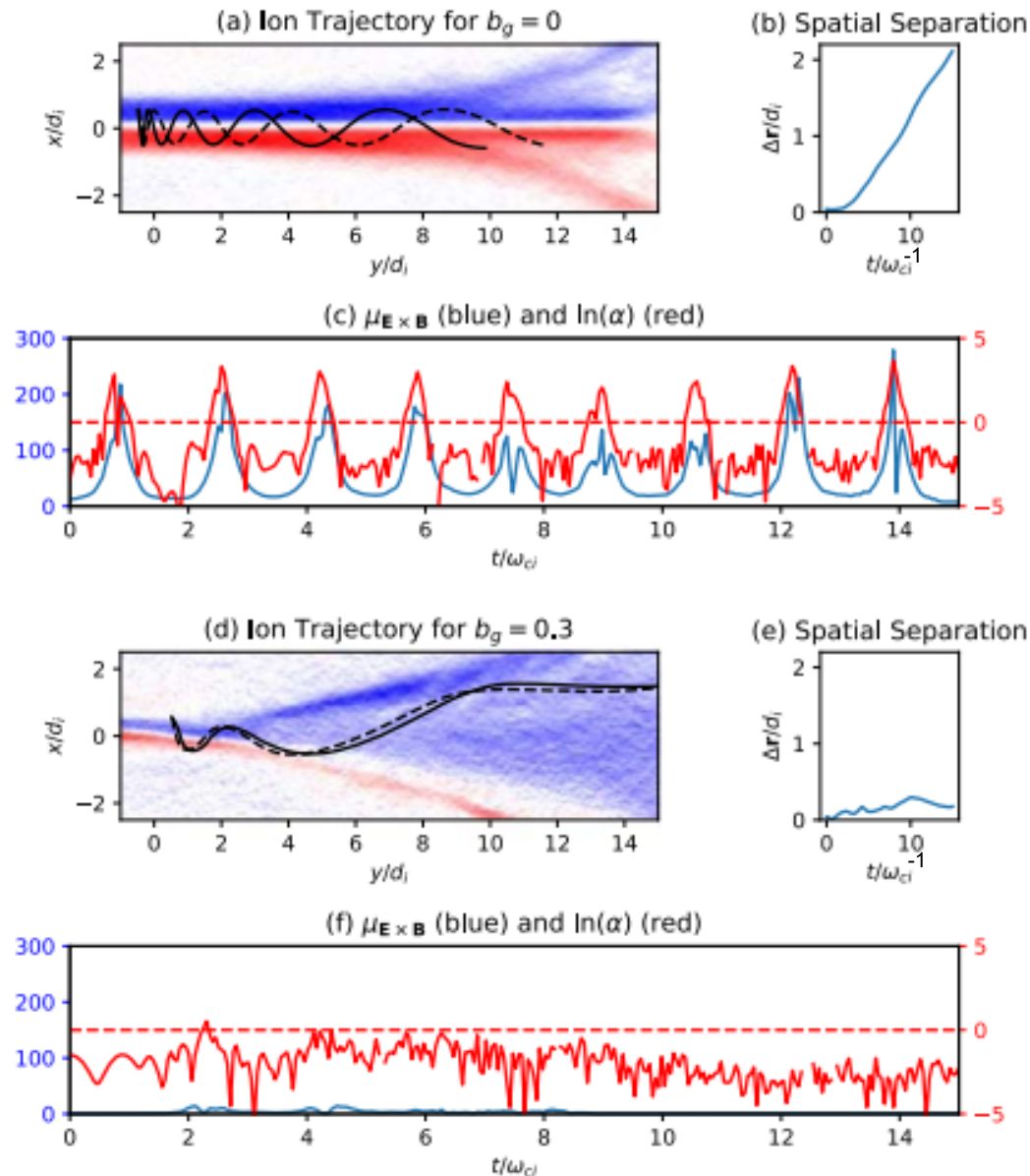
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Test-Particle Simulation



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Heavy ions satisfy the criterion more easily!

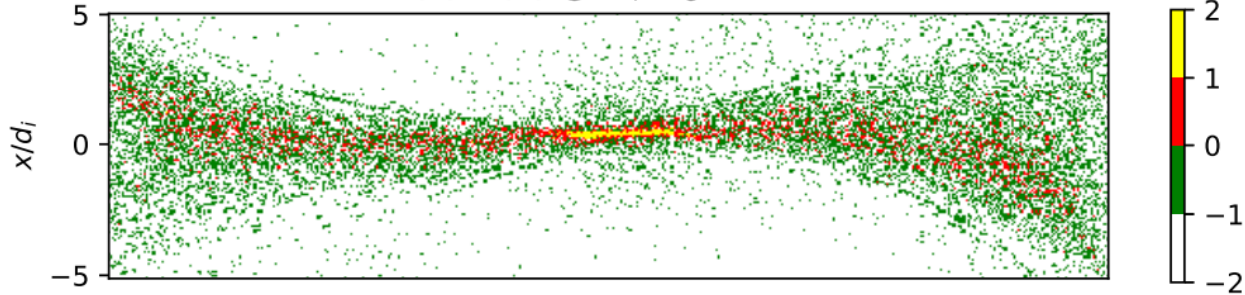
$$\frac{m_i}{q_i B^2} |\nabla_{\perp}^2 \phi| > 1$$

Heavy Ions

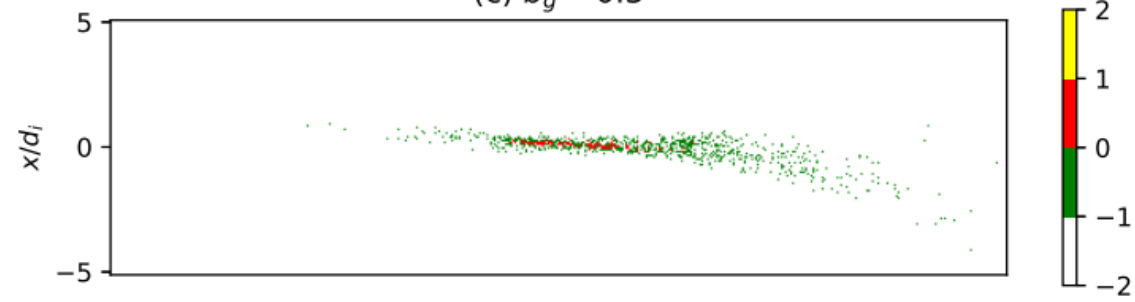
$$m_i/m_e = 500$$

$$m_i/m_e = 100$$

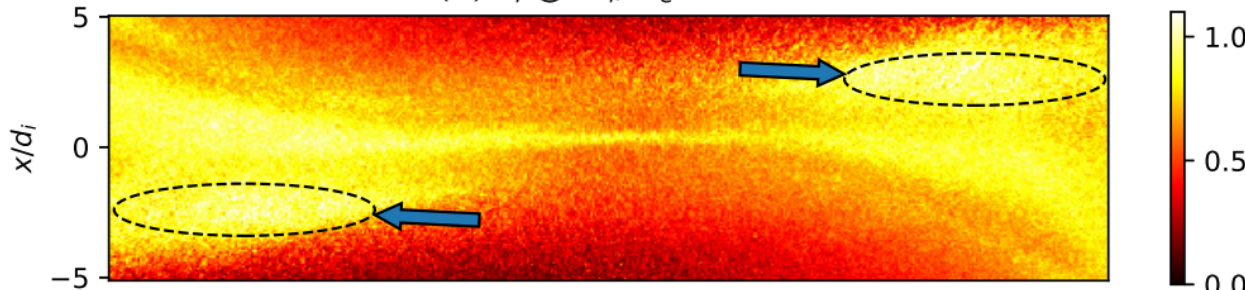
(a) $\ln(\alpha)$ @ $m_i/m_e = 500$



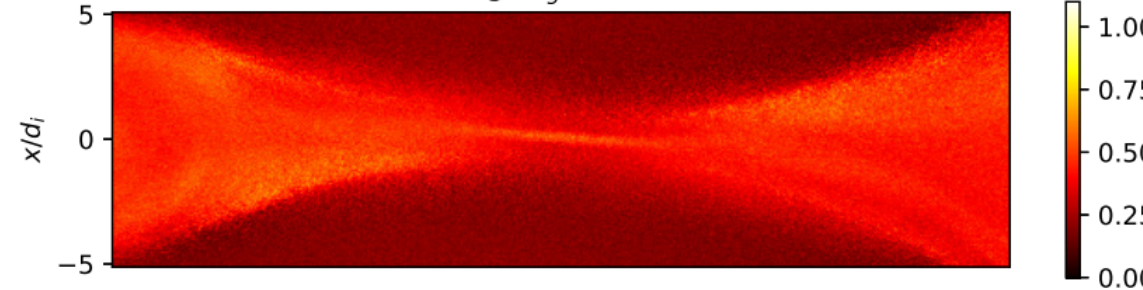
(c) $b_g = 0.3$



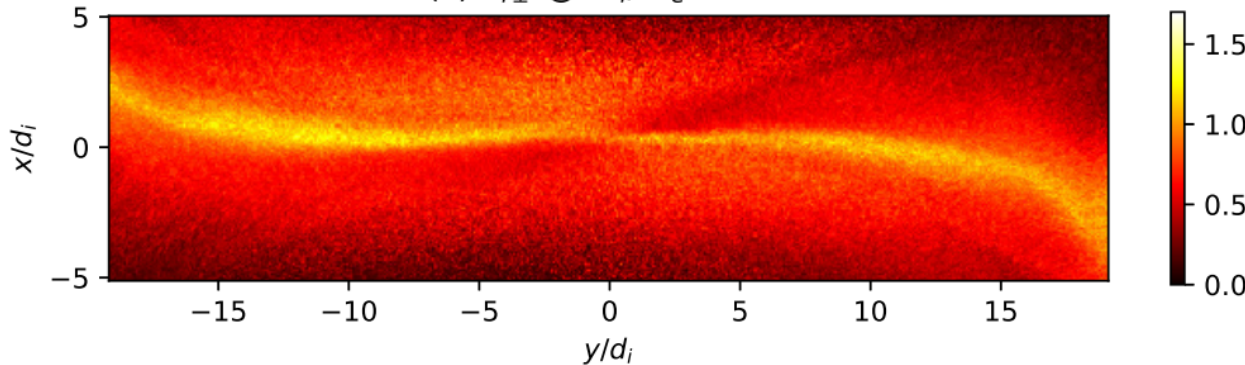
(b) T_i @ $m_i/m_e = 500$



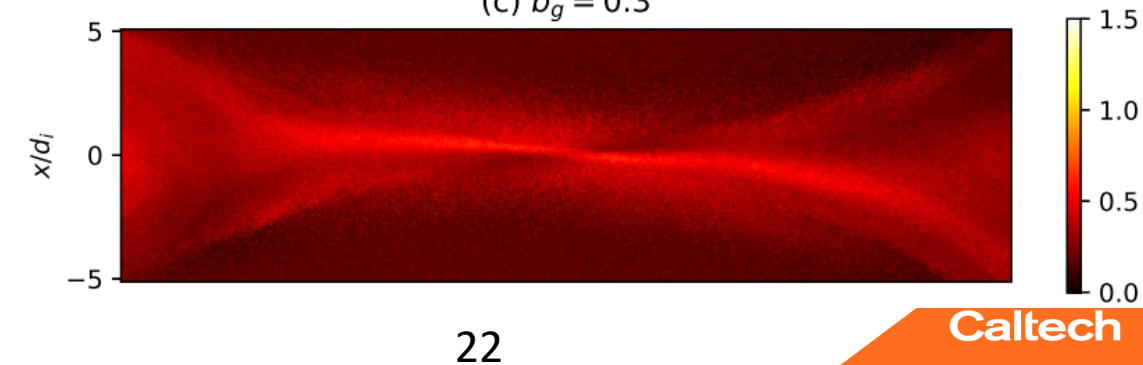
(g) $b_g = 0.3$



(c) $T_{i\perp}$ @ $m_i/m_e = 500$



(c) $b_g = 0.3$



Conclusion

- Stochastic ion heating is intrinsic to collisionless magnetic reconnection up to moderate guide fields
- In-plane Hall electric fields render ion motions chaotic and heat the ions
- Supported by fluid analysis, kinetic analysis, and particle-in-cell simulations

Thank you Questions?

Y. D. Yoon and P. M. Bellan, ApJL, 868, L31 (2018)

Y. D. Yoon and P. M. Bellan, ApJL, 887, L29 (2019)